

# $\rho(770)$

$$J^{PC} = 1^{+}(1^{-}-)$$

## THE $\rho(770)$

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The determination of the parameters of the  $\rho(770)$  is beset with many difficulties because of its large width. In physical region fits, the line shape does not correspond to a relativistic Breit-Wigner function with a  $P$ -wave width, but requires some additional shape parameter. This dependence on parameterization was demonstrated long ago [1]. Bose-Einstein correlations are another source of shifts in the  $\rho(770)$  line shape, particularly in multiparticle final-state systems [2].

The same model dependence afflicts any other source of resonance parameters, such as the energy dependence of the phase shift  $\delta_1^1$ , or the pole position. It is, therefore, not surprising that a study of  $\rho(770)$  dominance in the decays of the  $\eta$  and  $\eta'$  reveals the need for specific dynamical effects, in addition to the  $\rho(770)$  pole [3,4].

The cleanest determination of the  $\rho(770)$  mass and width comes from  $e^+e^-$  annihilation and  $\tau$ -lepton decays. Analysis of ALEPH [5] showed that the charged  $\rho(770)$  parameters measured from  $\tau$ -lepton decays are consistent with those of the neutral one determined from  $e^+e^-$  data [6]. This conclusion is qualitatively supported by the later studies of CLEO [7] and Belle [8]. However, comparison of the two-pion mass spectrum in  $\tau$  decays from OPAL [9], CLEO [7], and ALEPH [10,11], and the  $e^+e^- \rightarrow \pi^+\pi^-$  cross section from CMD-2 [12,13], showed significant discrepancies between the two shapes which can be as high as 10% above the  $\rho$  meson [14,15]. This discrepancy remains after measurements of the two-pion cross section in  $e^+e^-$

annihilation at KLOE [16,17,18,19], SND [20,21], BaBar [22] and, more recently BESIII [23] The effect is not accounted for by isospin breaking [24,25,26,27], but the accuracy of its calculation may be overestimated [28,29].

This problem seems to be solved after a recent analysis in [30] which showed that after correcting the  $\tau$  data for the missing  $\rho - \gamma$  mixing contribution, besides the other known isospin symmetry violating corrections, the  $\pi\pi$  I=1 part of the hadronic vacuum polarization contribution to the muon  $g - 2$  is fully compatible between  $\tau$  based and  $e^+e^-$  based evaluations. The global fit of the whole set of the  $\rho$ ,  $\omega$ , and  $\phi$  decays, taking into account mixing effects in the hidden local symmetry model, also showed consistency of the data on  $\tau$  decays to two pions and  $e^+e^-$  annihilation [31,32]. However, because of the progress in  $e^+e^-$  data, the  $\tau$  input is now less precise and less reliable due to additional theoretical uncertainties [33] decreasing importance of  $\tau$  versus  $e^+e^-$  comparison for the determination of  $\rho(770)$  parameters and other applications, like, e.g., calculations of hadronic vacuum polarization.

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### $\rho(770)$ MASS

We no longer list *S*-wave Breit-Wigner fits, or data with high combinatorial background.

**NEUTRAL ONLY,  $e^+e^-$** 

<u>VALUE (MeV)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>775.26±0.25 OUR AVERAGE</b>				
775.02±0.35		<sup>1</sup> LEES 12G	BABR	$e^+e^- \rightarrow \pi^+\pi^-\gamma$
775.97±0.46±0.70	900k	<sup>2</sup> AKHMETSHIN 07		$e^+e^- \rightarrow \pi^+\pi^-$
774.6 ±0.4 ±0.5	800k	<sup>3,4</sup> ACHASOV 06	SND	$e^+e^- \rightarrow \pi^+\pi^-$
775.65±0.64±0.50	114k	<sup>5,6</sup> AKHMETSHIN 04	CMD2	$e^+e^- \rightarrow \pi^+\pi^-$
775.9 ±0.5 ±0.5	1.98M	<sup>7</sup> ALOISIO 03	KLOE	1.02 $e^+e^- \rightarrow \pi^+\pi^-\pi^0$
775.8 ±0.9 ±2.0	500k	<sup>7</sup> ACHASOV 02	SND	1.02 $e^+e^- \rightarrow \pi^+\pi^-\pi^0$
775.9 ±1.1		<sup>8</sup> BARKOV 85	OLYA	$e^+e^- \rightarrow \pi^+\pi^-$
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
763.49±0.53		<sup>9</sup> BARTOS 17	RVUE	$e^+e^- \rightarrow \pi^+\pi^-$
758.23±0.46		<sup>10</sup> BARTOS 17A	RVUE	$e^+e^- \rightarrow \pi^+\pi^-$
775.8 ±0.5 ±0.3	1.98M	<sup>11</sup> ALOISIO 03	KLOE	1.02 $e^+e^- \rightarrow \pi^+\pi^-\pi^0$
775.9 ±0.6 ±0.5	1.98M	<sup>12</sup> ALOISIO 03	KLOE	1.02 $e^+e^- \rightarrow \pi^+\pi^-\pi^0$
775.0 ±0.6 ±1.1	500k	<sup>13</sup> ACHASOV 02	SND	1.02 $e^+e^- \rightarrow \pi^+\pi^-\pi^0$
775.1 ±0.7 ±5.3		<sup>14</sup> BENAYOUN 98	RVUE	$e^+e^- \rightarrow \pi^+\pi^-$ ,
				$\mu^+\mu^-$
770.5 ±1.9 ±5.1		<sup>15</sup> GARDNER 98	RVUE	0.28–0.92 $e^+e^- \rightarrow$
				$\pi^+\pi^-$
764.1 ±0.7		<sup>16</sup> O'CONNELL 97	RVUE	$e^+e^- \rightarrow \pi^+\pi^-$
757.5 ±1.5		<sup>17</sup> BERNICHA 94	RVUE	$e^+e^- \rightarrow \pi^+\pi^-$
768 ±1		<sup>18</sup> GESHKEN... 89	RVUE	$e^+e^- \rightarrow \pi^+\pi^-$

<sup>1</sup> Using the GOUNARIS 68 parametrization with the complex phase of the  $\rho$ - $\omega$  interference and leaving the masses and widths of the  $\rho(1450)$ ,  $\rho(1700)$ , and  $\rho(2150)$  resonances as free parameters of the fit.

<sup>2</sup> A combined fit of AKHMETSHIN 07, AULCHENKO 06, and AULCHENKO 05.

<sup>3</sup> Supersedes ACHASOV 05A.

<sup>4</sup> A fit of the SND data from 400 to 1000 MeV using parameters of the  $\rho(1450)$  and  $\rho(1700)$  from a fit of the data of BARKOV 85, BISELLO 89 and ANDERSON 00A.

<sup>5</sup> Using the GOUNARIS 68 parametrization with the complex phase of the  $\rho$ - $\omega$  interference.

<sup>6</sup> Update of AKHMETSHIN 02.

<sup>7</sup> Assuming  $m_{\rho^+} = m_{\rho^-}$ ,  $\Gamma_{\rho^+} = \Gamma_{\rho^-}$ .

<sup>8</sup> From the GOUNARIS 68 parametrization of the pion form factor.

<sup>9</sup> Applies the Unitary & Analytic Model of the pion electromagnetic form factor of DUBNICKA 10 to analyze the data of LEES 12G and ABLIKIM 16C.

<sup>10</sup> Applies the Unitary & Analytic Model of the pion electromagnetic form factor of DUBNICKA 10 to analyze the data of ACHASOV 06, AKHMETSHIN 07, AUBERT 09AS, and AMBROSINO 11A.

<sup>11</sup> Assuming  $m_{\rho^+} = m_{\rho^-} = m_{\rho^0}$ ,  $\Gamma_{\rho^+} = \Gamma_{\rho^-} = \Gamma_{\rho^0}$ .

<sup>12</sup> Without limitations on masses and widths.

<sup>13</sup> Assuming  $m_{\rho^0} = m_{\rho^\pm}$ ,  $g_{\rho^0\pi\pi} = g_{\rho^\pm\pi\pi}$ .

<sup>14</sup> Using the data of BARKOV 85 in the hidden local symmetry model.

<sup>15</sup> From the fit to  $e^+e^- \rightarrow \pi^+\pi^-$  data from the compilations of HEYN 81 and BARKOV 85, including the GOUNARIS 68 parametrization of the pion form factor.

<sup>16</sup> A fit of BARKOV 85 data assuming the direct  $\omega\pi\pi$  coupling.

<sup>17</sup> Applying the S-matrix formalism to the BARKOV 85 data.

<sup>18</sup> Includes BARKOV 85 data. Model-dependent width definition.

**CHARGED ONLY,  $\tau$  DECAYS and  $e^+e^-$** 

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>775.11±0.34 OUR AVERAGE</b>					
774.6 ±0.2 ±0.5	5.4M	<sup>1,2</sup> FUJIKAWA	08	BELL ±	$\tau^- \rightarrow \pi^- \pi^0 \nu_\tau$
775.5 ±0.7		<sup>2,3</sup> SCHAELE	05c	ALEP	$\tau^- \rightarrow \pi^- \pi^0 \nu_\tau$
775.5 ±0.5 ±0.4	1.98M	<sup>4</sup> ALOISIO	03	KLOE	$1.02 e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$
775.1 ±1.1 ±0.5	87k	<sup>5,6</sup> ANDERSON	00A	CLE2	$\tau^- \rightarrow \pi^- \pi^0 \nu_\tau$
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●					
761.60±0.95		<sup>7</sup> BARTOS	17A	RVUE	$\tau^- \rightarrow \pi^- \pi^0 \nu_\tau$
774.8 ±0.6 ±0.4	1.98M	<sup>8</sup> ALOISIO	03	KLOE -	$1.02 e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$
776.3 ±0.6 ±0.7	1.98M	<sup>8</sup> ALOISIO	03	KLOE +	$1.02 e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$
773.9 ±2.0 <sup>+0.3</sup> <sub>-1.0</sub>		<sup>9</sup> SANZ-CILLERO3		RVUE	$\tau^- \rightarrow \pi^- \pi^0 \nu_\tau$
774.5 ±0.7 ±1.5	500k	<sup>4</sup> ACHASOV	02	SND ±	$1.02 e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$
775.1 ±0.5		<sup>10</sup> PICH	01	RVUE	$\tau^- \rightarrow \pi^- \pi^0 \nu_\tau$

<sup>1</sup>  $|F_\pi(0)|^2$  fixed to 1.<sup>2</sup> From the GOUNARIS 68 parametrization of the pion form factor.<sup>3</sup> The error combines statistical and systematic uncertainties. Supersedes BARATE 97M.<sup>4</sup> Assuming  $m_{\rho^+} = m_{\rho^-}$ ,  $\Gamma_{\rho^+} = \Gamma_{\rho^-}$ .<sup>5</sup>  $\rho(1700)$  mass and width fixed at 1700 MeV and 235 MeV respectively.<sup>6</sup> From the GOUNARIS 68 parametrization of the pion form factor. The second error is a model error taking into account different parametrizations of the pion form factor.<sup>7</sup> Applies the Unitary & Analytic Model of the pion electromagnetic form factor of DUBNICKA 10 to analyze the data of FUJIKAWA 08.<sup>8</sup> Without limitations on masses and widths.<sup>9</sup> Using the data of BARATE 97M and the effective chiral Lagrangian.<sup>10</sup> From a fit of the model-independent parameterization of the pion form factor to the data of BARATE 97M.**MIXED CHARGES, OTHER REACTIONS**

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>763.0±0.3±1.2</b>	600k	<sup>1</sup> ABELE	99E	CBAR	$0 \pm 0.0 \bar{p} p \rightarrow \pi^+ \pi^- \pi^0$

<sup>1</sup> Assuming the equality of  $\rho^+$  and  $\rho^-$  masses and widths.**CHARGED ONLY, HADROPRODUCED**

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>766.5±1.1 OUR AVERAGE</b>					
763.7±3.2		ABELE	97	CBAR	$\bar{p} n \rightarrow \pi^- \pi^0 \pi^0$
768 ±9		AGUILAR-...	91	EHS	400 $pp$
767 ±3	2935	<sup>1</sup> CAPRARO	87	SPEC -	200 $\pi^- \text{Cu} \rightarrow \pi^- \pi^0 \text{Cu}$
761 ±5	967	<sup>1</sup> CAPRARO	87	SPEC -	200 $\pi^- \text{Pb} \rightarrow \pi^- \pi^0 \text{Pb}$
771 ±4		HUSTON	86	SPEC +	202 $\pi^+ \text{A} \rightarrow \pi^+ \pi^0 \text{A}$
766 ±7	6500	<sup>2</sup> BYERLY	73	OSPK -	5 $\pi^- p$
766.8±1.5	9650	<sup>3</sup> PISUT	68	RVUE -	1.7–3.2 $\pi^- p$ , $t < 10$
767 ±6	900	<sup>1</sup> EISNER	67	HBC -	4.2 $\pi^- p$ , $t < 10$

<sup>1</sup> Mass errors enlarged by us to  $\Gamma/\sqrt{N}$ ; see the note with the  $K^*(892)$  mass.<sup>2</sup> Phase shift analysis. Systematic errors added corresponding to spread of different fits.<sup>3</sup> From fit of 3-parameter relativistic  $P$ -wave Breit-Wigner to total mass distribution. Includes BATON 68, MILLER 67B, ALFF-STEINBERGER 66, HAGOPIAN 66, HAGOPIAN 66B, JACOBS 66B, JAMES 66, WEST 66, BLIEDEN 65 and CARMONY 64.

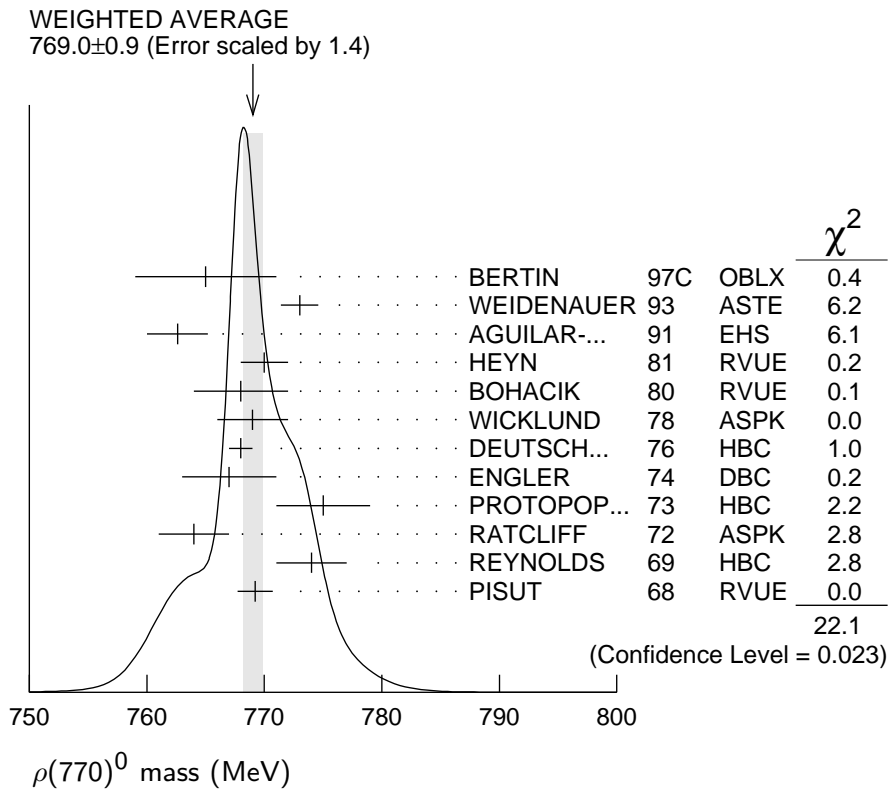
**NEUTRAL ONLY, PHOTOPRODUCED**

<u>VALUE (MeV)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>769.0 ± 1.0 OUR AVERAGE</b>				
771 ± 2 $\begin{smallmatrix} +2 \\ -1 \end{smallmatrix}$	63.5k	<sup>1</sup> ABRAMOWICZ12	ZEUS	$ep \rightarrow e\pi^+\pi^-p$
770 ± 2 ±1	79k	<sup>2</sup> BREITWEG	98B ZEUS	50–100 $\gamma p$
767.6 ± 2.7		BARTALUCCI	78 CNTR	$\gamma p \rightarrow e^+e^-p$
775 ± 5		GLADDING	73 CNTR	2.9–4.7 $\gamma p$
767 ± 4	1930	BALLAM	72 HBC	2.8 $\gamma p$
770 ± 4	2430	BALLAM	72 HBC	4.7 $\gamma p$
765 ± 10		ALVENSLEB...	70 CNTR	$\gamma A, t < 0.01$
767.7 ± 1.9	140k	BIGGS	70 CNTR	$< 4.1 \gamma C \rightarrow \pi^+\pi^-C$
765 ± 5	4000	ASBURY	67B CNTR	$\gamma + Pb$
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
771 ± 2	79k	<sup>3</sup> BREITWEG	98B ZEUS	50–100 $\gamma p$

<sup>1</sup> Using the KUHN 90 parametrization of the pion form factor, neglecting  $\rho$ – $\omega$  interference.<sup>2</sup> From the parametrization according to SOEDING 66.<sup>3</sup> From the parametrization according to ROSS 66.**NEUTRAL ONLY, OTHER REACTIONS**

<u>VALUE (MeV)</u>	<u>EVTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
<b>769.0 ± 0.9 OUR AVERAGE</b>		Error includes scale factor of 1.4.		See the ideogram below.
765 ± 6		BERTIN	97C OBLX	0.0 $\bar{p}p \rightarrow \pi^+\pi^-\pi^0$
773 ± 1.6		WEIDENAUER	93 ASTE	$\bar{p}p \rightarrow \pi^+\pi^- \omega$
762.6 ± 2.6		AGUILAR-...	91 EHS	400 $pp$
770 ± 2		<sup>1</sup> HEYN	81 RVUE	Pion form factor
768 ± 4		<sup>2,3</sup> BOHACIK	80 RVUE	
769 ± 3		<sup>4</sup> WICKLUND	78 ASPK	3,4,6 $\pi^\pm N$
768 ± 1	76k	DEUTSCH...	76 HBC	16 $\pi^+ p$
767 ± 4	4100	ENGLER	74 DBC	6 $\pi^+ n \rightarrow \pi^+\pi^- p$
775 ± 4	32k	<sup>2</sup> PROTOPOP...	73 HBC	7.1 $\pi^+ p, t < 0.4$
764 ± 3	6.8k	<sup>5</sup> RATCLIFF	72 ASPK	15 $\pi^- p, t < 0.3$
774 ± 3	1.7k	REYNOLDS	69 HBC	2.26 $\pi^- p$
769.2 ± 1.5	13.3k	<sup>6</sup> PISUT	68 RVUE	1.7–3.2 $\pi^- p, t < 10$
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
774.34 ± 0.18 ± 0.35	970k	<sup>7</sup> ABLIKIM	18C BES3	$\eta'(958) \rightarrow \gamma\pi^+\pi^-$
772.93 ± 0.18 ± 0.34	970k	<sup>8</sup> ABLIKIM	18C BES3	$\eta'(958) \rightarrow \gamma\pi^+\pi^-$
773.5 ± 2.5		<sup>9</sup> COLANGELO	01 RVUE	$\pi\pi \rightarrow \pi\pi$
762.3 ± 0.5 ± 1.2	600k	<sup>10</sup> ABELE	99E CBAR	0.0 $\bar{p}p \rightarrow \pi^+\pi^-\pi^0$
777 ± 2	4.9k	<sup>11</sup> ADAMS	97 E665	470 $\mu p \rightarrow \mu XB$
770 ± 2		<sup>12</sup> BOGOLYUB...	97 MIRA	32 $\bar{p}p \rightarrow \pi^+\pi^- X$
768 ± 8		<sup>12</sup> BOGOLYUB...	97 MIRA	32 $pp \rightarrow \pi^+\pi^- X$
761.1 ± 2.9		DUBNICKA	89 RVUE	$\pi$ form factor
777.4 ± 2.0		<sup>13</sup> CHABAUD	83 ASPK	17 $\pi^- p$ polarized
769.5 ± 0.7		<sup>2,3</sup> LANG	79 RVUE	
770 ± 9		<sup>3</sup> ESTABROOKS	74 RVUE	17 $\pi^- p \rightarrow \pi^+\pi^- n$
773.5 ± 1.7	11.2k	<sup>14</sup> JACOBS	72 HBC	2.8 $\pi^- p$
775 ± 3	2.2k	<sup>15</sup> HYAMS	68 OSPK	11.2 $\pi^- p$

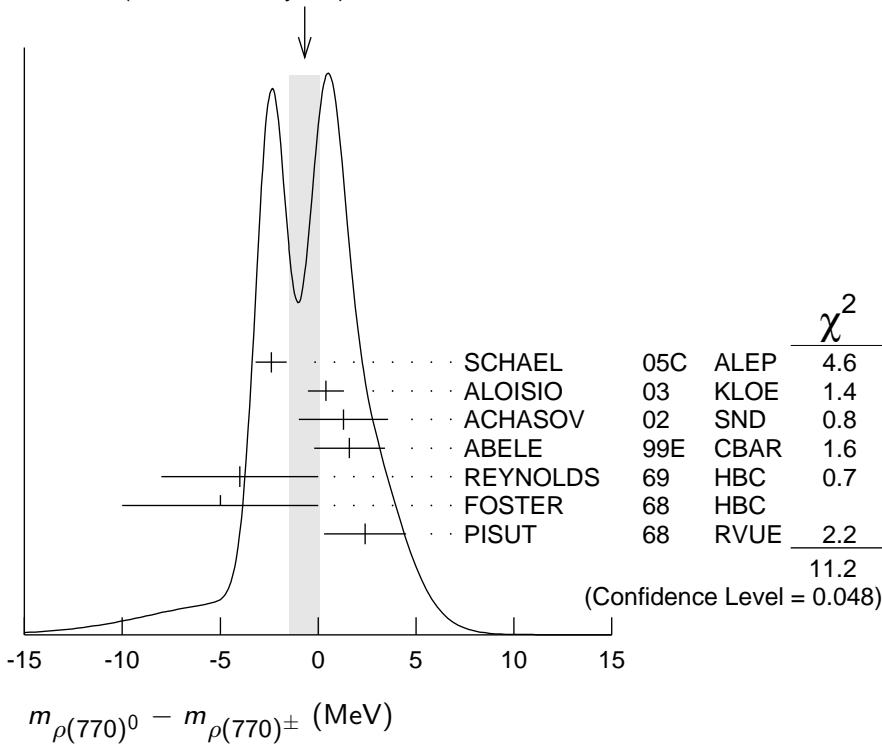
- <sup>1</sup> HEYN 81 includes all spacelike and timelike  $F_\pi$  values until 1978.
- <sup>2</sup> From pole extrapolation.
- <sup>3</sup> From phase shift analysis of GRAYER 74 data.
- <sup>4</sup> Phase shift analysis. Systematic errors added corresponding to spread of different fits.
- <sup>5</sup> Published values contain misprints. Corrected by private communication RATCLIFF 74.
- <sup>6</sup> Includes MALAMUD 69, ARMENISE 68, BACON 67, HUWE 67, MILLER 67B, ALFF-STEINBERGER 66, HAGOPIAN 66, HAGOPIAN 66B, JACOBS 66B, JAMES 66, WEST 66, GOLDHABER 64, ABOLINS 63.
- <sup>7</sup> From a fit to  $\pi^+\pi^-$  mass using  $\rho(770)$  (parametrized with the Gounaris-Sakurai approach),  $\omega(782)$ , and box anomaly components.
- <sup>8</sup> From a fit to  $\pi^+\pi^-$  mass using  $\rho(770)$  (parametrized with the Gounaris-Sakurai approach),  $\omega(782)$ , and  $\rho(1450)$  components.
- <sup>9</sup> Breit-Wigner mass from a phase-shift analysis of HYAMS 73 and PROTOPOPESCU 73 data.
- <sup>10</sup> Using relativistic Breit-Wigner and taking into account  $\rho$ - $\omega$  interference.
- <sup>11</sup> Systematic errors not evaluated.
- <sup>12</sup> Systematic effects not studied.
- <sup>13</sup> From fit of 3-parameter relativistic Breit-Wigner to helicity-zero part of P-wave intensity. CHABAUD 83 includes data of GRAYER 74.
- <sup>14</sup> Mass errors enlarged by us to  $\Gamma/\sqrt{N}$ ; see the note with the  $K^*(892)$  mass.
- <sup>15</sup> Of HYAMS 68 six parametrizations, this is theoretically soundest. MR



**$m_{\rho(770)^0} - m_{\rho(770)^\pm}$**

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>-0.7 ±0.8</b>	<b>OUR AVERAGE</b>	Error includes scale factor of 1.5. See the ideogram below.			
-2.4 ±0.8		<sup>1</sup> SCHAEL	05C	ALEP	$\tau^- \rightarrow \pi^- \pi^0 \nu_\tau$
0.4 ±0.7 ±0.6	1.98M	<sup>2</sup> ALOISIO	03	KLOE	$1.02 e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$
1.3 ±1.1 ±2.0	500k	<sup>2</sup> ACHASOV	02	SND	$1.02 e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$
1.6 ±0.6 ±1.7	600k	ABELE	99E	CBAR	$\pm 0.0 \bar{p} p \rightarrow \pi^+ \pi^- \pi^0$
-4 ±4	3000	<sup>3</sup> REYNOLDS	69	HBC	-0 2.26 $\pi^- p$
-5 ±5	3600	<sup>3</sup> FOSTER	68	HBC	$\pm 0.0 \bar{p} p$
2.4 ±2.1	22950	<sup>4</sup> PISUT	68	RVUE	$\pi N \rightarrow \rho N$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
-3.37 ±1.06		<sup>5</sup> BARTOS	17A	RVUE	$e^+ e^- \rightarrow \pi^+ \pi^-$ , $\tau^- \rightarrow \pi^- \pi^0 \nu_\tau$

WEIGHTED AVERAGE  
-0.7±0.8 (Error scaled by 1.5)



<sup>1</sup> From the combined fit of the  $\tau^-$  data from ANDERSON 00A and SCHAEL 05C and  $e^+ e^-$  data from the compilation of BARKOV 85, AKHMETSHIN 04, and ALOISIO 05. Supersedes BARATE 97M.

<sup>2</sup> Assuming  $m_{\rho^+} = m_{\rho^-}$ ,  $\Gamma_{\rho^+} = \Gamma_{\rho^-}$ .

<sup>3</sup> From quoted masses of charged and neutral modes.

<sup>4</sup> Includes MALAMUD 69, ARMENISE 68, BATON 68, BACON 67, HUWE 67, MILLER 67B, ALFF-STEINBERGER 66, HAGOPIAN 66, HAGOPIAN 66B, JACOBS 66B, JAMES 66, WEST 66, BLIEDEN 65, CARMONY 64, GOLDHABER 64, ABOLINS 63.



<sup>5</sup>Applies the Unitary & Analytic Model of the pion electromagnetic form factor of DUBNICKA 10 to analyze the data of ACHASOV 06, AKHMETSHIN 07, AUBERT 09AS, AMBROSINO 11A, and FUJIKAWA 08.

### $m_{\rho(770)^+} - m_{\rho(770)^-}$

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
$1.5 \pm 0.8 \pm 0.7$	1.98M	<sup>1</sup> ALOISIO	03	KLOE 1.02 $e^+e^- \rightarrow \pi^+\pi^-\pi^0$
<sup>1</sup> Without limitations on masses and widths.				

### $\rho(770)$ RANGE PARAMETER

The range parameter  $R$  enters an energy-dependent correction to the width, of the form  $(1 + q_r^2 R^2) / (1 + q^2 R^2)$ , where  $q$  is the momentum of one of the pions in the  $\pi\pi$  rest system. At resonance,  $q = q_r$ .

VALUE ( $\text{GeV}^{-1}$ )	DOCUMENT ID	TECN	CHG	COMMENT
$5.3^{+0.9}_{-0.7}$	<sup>1</sup> CHABAUD	83	ASPK	0 17 $\pi^- p$ polarized
<sup>1</sup> The old PISUT 68 value, properly corrected, was $3.2 \pm 0.6$ .				

### $\rho(770)$ WIDTH

We no longer list  $S$ -wave Breit-Wigner fits, or data with high combinatorial background.

#### NEUTRAL ONLY, $e^+e^-$

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b><math>147.8 \pm 0.9</math></b>	<b>OUR AVERAGE</b>	Error includes scale factor of 2.0.		See the ideogram below.
$149.59 \pm 0.67$		<sup>1</sup> LEES	12G	BABR $e^+e^- \rightarrow \pi^+\pi^-\gamma$
$145.98 \pm 0.75 \pm 0.50$	900k	<sup>2</sup> AKHMETSHIN 07		$e^+e^- \rightarrow \pi^+\pi^-$
$146.1 \pm 0.8 \pm 1.5$	800k	<sup>3,4</sup> ACHASOV 06	SND	$e^+e^- \rightarrow \pi^+\pi^-$
$143.85 \pm 1.33 \pm 0.80$	114k	<sup>5,6</sup> AKHMETSHIN 04	CMD2	$e^+e^- \rightarrow \pi^+\pi^-$
$147.3 \pm 1.5 \pm 0.7$	1.98M	<sup>7</sup> ALOISIO 03	KLOE	1.02 $e^+e^- \rightarrow \pi^+\pi^-\pi^0$
$151.1 \pm 2.6 \pm 3.0$	500k	<sup>7</sup> ACHASOV 02	SND	1.02 $e^+e^- \rightarrow \pi^+\pi^-\pi^0$
$150.5 \pm 3.0$		<sup>8</sup> BARKOV 85	OLYA	$e^+e^- \rightarrow \pi^+\pi^-$
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
$144.06 \pm 0.85$		<sup>9</sup> BARTOS 17	RVUE	$e^+e^- \rightarrow \pi^+\pi^-$
$144.56 \pm 0.80$		<sup>10</sup> BARTOS 17A	RVUE	$e^+e^- \rightarrow \pi^+\pi^-$
$143.9 \pm 1.3 \pm 1.1$	1.98M	<sup>11</sup> ALOISIO 03	KLOE	1.02 $e^+e^- \rightarrow \pi^+\pi^-\pi^0$
$147.4 \pm 1.5 \pm 0.7$	1.98M	<sup>12</sup> ALOISIO 03	KLOE	1.02 $e^+e^- \rightarrow \pi^+\pi^-\pi^0$
$149.8 \pm 2.2 \pm 2.0$	500k	<sup>13</sup> ACHASOV 02	SND	1.02 $e^+e^- \rightarrow \pi^+\pi^-\pi^0$

147.9 ±1.5 ±7.5	14	BENAYOUN	98	RVUE	$e^+e^- \rightarrow \pi^+\pi^-,$ $\mu^+\mu^-$
153.5 ±1.3 ±4.6	15	GARDNER	98	RVUE	0.28–0.92 $e^+e^- \rightarrow$ $\pi^+\pi^-$
145.0 ±1.7	16	O'CONNELL	97	RVUE	$e^+e^- \rightarrow \pi^+\pi^-$
142.5 ±3.5	17	BERNICHIA	94	RVUE	$e^+e^- \rightarrow \pi^+\pi^-$
138 ±1	18	GESHKEN...	89	RVUE	$e^+e^- \rightarrow \pi^+\pi^-$

<sup>1</sup> Using the GOUNARIS 68 parametrization with the complex phase of the  $\rho$ - $\omega$  interference and leaving the masses and widths of the  $\rho(1450)$ ,  $\rho(1700)$ , and  $\rho(2150)$  resonances as free parameters of the fit.

<sup>2</sup> A combined fit of AKHMETSHIN 07, AULCHENKO 06, and AULCHENKO 05.

<sup>3</sup> Supersedes ACHASOV 05A.

<sup>4</sup> A fit of the SND data from 400 to 1000 MeV using parameters of the  $\rho(1450)$  and  $\rho(1700)$  from a fit of the data of BARKOV 85, BISELLO 89 and ANDERSON 00A.

<sup>5</sup> Using the GOUNARIS 68 parametrization with the complex phase of the  $\rho$ - $\omega$  interference.

<sup>6</sup> From a fit in the energy range 0.61 to 0.96 GeV. Update of AKHMETSHIN 02.

<sup>7</sup> Assuming  $m_{\rho^+} = m_{\rho^-}$ ,  $\Gamma_{\rho^+} = \Gamma_{\rho^-}$ .

<sup>8</sup> From the GOUNARIS 68 parametrization of the pion form factor.

<sup>9</sup> Applies the Unitary & Analytic Model of the pion electromagnetic form factor of DUBNICKA 10 to analyze the data of LEES 12G and ABLIKIM 16C.

<sup>10</sup> Applies the Unitary & Analytic Model of the pion electromagnetic form factor of DUBNICKA 10 to analyze the data of ACHASOV 06, AKHMETSHIN 07, AUBERT 09AS, and AMBROSINO 11A.

<sup>11</sup> Assuming  $m_{\rho^+} = m_{\rho^-} = m_{\rho^0}$ ,  $\Gamma_{\rho^+} = \Gamma_{\rho^-} = \Gamma_{\rho^0}$ .

<sup>12</sup> Without limitations on masses and widths.

<sup>13</sup> Assuming  $m_{\rho^0} = m_{\rho^\pm}$ ,  $g_{\rho^0\pi\pi} = g_{\rho^\pm\pi\pi}$ .

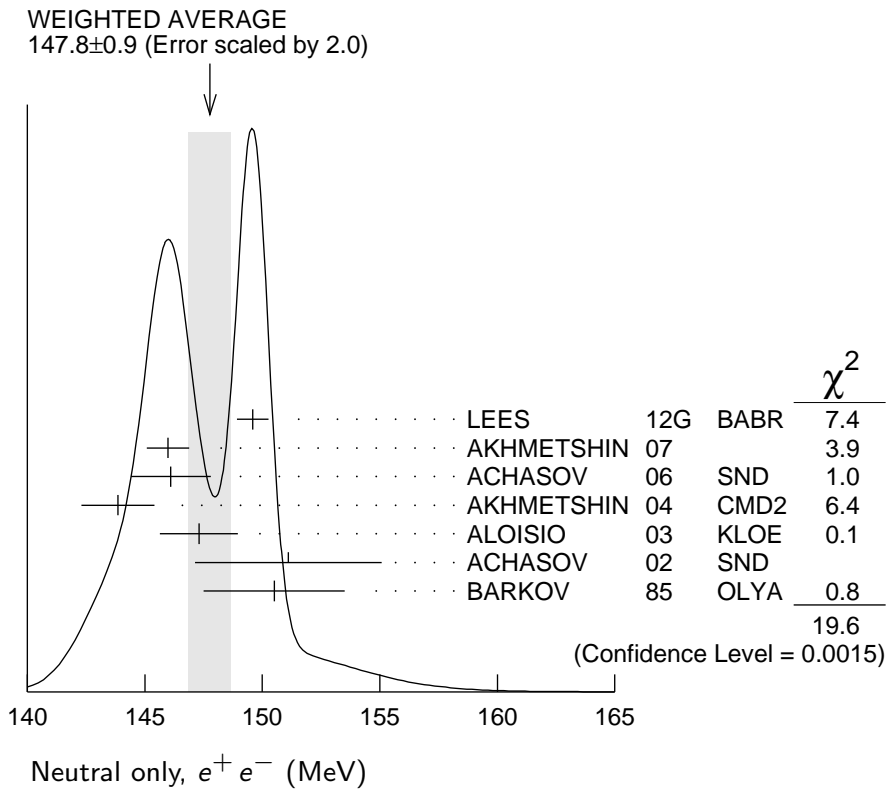
<sup>14</sup> Using the data of BARKOV 85 in the hidden local symmetry model.

<sup>15</sup> From the fit to  $e^+e^- \rightarrow \pi^+\pi^-$  data from the compilations of HEYN 81 and BARKOV 85, including the GOUNARIS 68 parametrization of the pion form factor.

<sup>16</sup> A fit of BARKOV 85 data assuming the direct  $\omega\pi\pi$  coupling.

<sup>17</sup> Applying the S-matrix formalism to the BARKOV 85 data.

<sup>18</sup> Includes BARKOV 85 data. Model-dependent width definition.



### CHARGED ONLY, $\tau$ DECAYS and $e^+e^-$

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>149.1 ±0.8</b>					<b>OUR FIT</b>
<b>149.1 ±0.8</b>					<b>OUR AVERAGE</b>
148.1 ±0.4 ±1.7	5.4M	<sup>1,2</sup> FUJIKAWA	08	BELL	± $\tau^- \rightarrow \pi^- \pi^0 \nu_\tau$
149.0 ±1.2		<sup>2,3</sup> SCHAEEL	05C	ALEP	$\tau^- \rightarrow \pi^- \pi^0 \nu_\tau$
149.9 ±2.3 ±2.0	500k	<sup>4</sup> ACHASOV	02	SND	± $1.02 e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$
150.4 ±1.4 ±1.4	87k	<sup>5,6</sup> ANDERSON	00A	CLE2	$\tau^- \rightarrow \pi^- \pi^0 \nu_\tau$
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●					
139.90±0.46		<sup>7</sup> BARTOS	17A	RVUE	$\tau^- \rightarrow \pi^- \pi^0 \nu_\tau$
143.7 ±1.3 ±1.2	1.98M	<sup>4</sup> ALOISIO	03	KLOE	± $1.02 e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$
142.9 ±1.3 ±1.4	1.98M	<sup>8</sup> ALOISIO	03	KLOE	- $1.02 e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$
144.7 ±1.4 ±1.2	1.98M	<sup>8</sup> ALOISIO	03	KLOE	+ $1.02 e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$
150.2 ±2.0 <sup>+0.7</sup> <sub>-1.6</sub>		<sup>9</sup> SANZ-CILLERO03		RVUE	$\tau^- \rightarrow \pi^- \pi^0 \nu_\tau$
150.9 ±2.2 ±2.0	500k	<sup>10</sup> ACHASOV	02	SND	$1.02 e^+ e^- \rightarrow \pi^+ \pi^- \pi^0$

<sup>1</sup>  $|F_\pi(0)|^2$  fixed to 1.

<sup>2</sup> From the GOUNARIS 68 parametrization of the pion form factor.

<sup>3</sup> The error combines statistical and systematic uncertainties. Supersedes BARATE 97M.

<sup>4</sup> Assuming  $m_{\rho^+} = m_{\rho^-}$ ,  $\Gamma_{\rho^+} = \Gamma_{\rho^-}$ .

<sup>5</sup>  $\rho(1700)$  mass and width fixed at 1700 MeV and 235 MeV respectively.

<sup>6</sup> From the GOUNARIS 68 parametrization of the pion form factor. The second error is a model error taking into account different parametrizations of the pion form factor.

<sup>7</sup> Applies the Unitary & Analytic Model of the pion electromagnetic form factor of DUB-NICKA 10 to analyze the data of FUJIKAWA 08.

<sup>8</sup> Without limitations on masses and widths.

<sup>9</sup> Using the data of BARATE 97M and the effective chiral Lagrangian.

<sup>10</sup> Assuming  $m_{\rho^0} = m_{\rho^\pm}$ ,  $g_{\rho^0\pi\pi} = g_{\rho^\pm\pi\pi}$ .

## MIXED CHARGES, OTHER REACTIONS

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>149.5±1.3</b>	600k	<sup>1</sup> ABELE	99E	CBAR	0± 0.0 $\bar{p}p \rightarrow \pi^+\pi^-\pi^0$

<sup>1</sup> Assuming the equality of  $\rho^+$  and  $\rho^-$  masses and widths.

## CHARGED ONLY, HADROPRODUCED

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>150.2± 2.4 OUR FIT</b>					
<b>150.2± 2.4 OUR AVERAGE</b>					
152.8± 4.3		ABELE	97	CBAR	$\bar{p}n \rightarrow \pi^-\pi^0\pi^0$
155 ±11	2.9k	<sup>1</sup> CAPRARO	87	SPEC	− 200 $\pi^- \text{Cu} \rightarrow \pi^-\pi^0\text{Cu}$
154 ±20	967	<sup>1</sup> CAPRARO	87	SPEC	− 200 $\pi^- \text{Pb} \rightarrow \pi^-\pi^0\text{Pb}$
150 ± 5		HUSTON	86	SPEC	+ 202 $\pi^+ \text{A} \rightarrow \pi^+\pi^0\text{A}$
146 ±12	6.5k	<sup>2</sup> BYERLY	73	OSPK	− 5 $\pi^- p$
148.2± 4.1	9.6k	<sup>3</sup> PISUT	68	RVUE	− 1.7–3.2 $\pi^- p$ , $t < 10$
146 ±13	900	EISNER	67	HBC	− 4.2 $\pi^- p$ , $t < 10$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
137.0± 0.4		<sup>4</sup> ABLIKIM	17	BES3	$J/\psi \rightarrow \gamma 3\pi$

<sup>1</sup> Width errors enlarged by us to  $4\Gamma/\sqrt{N}$ ; see the note with the  $K^*(892)$  mass.

<sup>2</sup> Phase shift analysis. Systematic errors added corresponding to spread of different fits.

<sup>3</sup> From fit of 3-parameter relativistic  $P$ -wave Breit-Wigner to total mass distribution. Includes BATON 68, MILLER 67B, ALFF-STEINBERGER 66, HAGOPIAN 66, HAGOPIAN 66B, JACOBS 66B, JAMES 66, WEST 66, BLIEDEN 65 and CARMONY 64.

<sup>4</sup>  $S$ -matrix pole at a fixed  $\rho$  meson mass of 775.49 MeV.

## NEUTRAL ONLY, PHOTOPRODUCED

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>151.7± 2.6 OUR AVERAGE</b>					
155 ± 5 ± 2	63.5k	<sup>1</sup> ABRAMOWICZ12	ZEUS		$ep \rightarrow e\pi^+\pi^-p$
146 ± 3 ±13	79k	<sup>2</sup> BREITWEG	98B	ZEUS	50–100 $\gamma p$
150.9± 3.0		BARTALUCCI	78	CNTR	$\gamma p \rightarrow e^+e^-p$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
138 ± 3	79k	<sup>3</sup> BREITWEG	98B	ZEUS	50–100 $\gamma p$
147 ±11		GLADDING	73	CNTR	2.9–4.7 $\gamma p$
155 ±12	2430	BALLAM	72	HBC	4.7 $\gamma p$
145 ±13	1930	BALLAM	72	HBC	2.8 $\gamma p$
140 ± 5		ALVENSLEB...	70	CNTR	$\gamma \text{A}$ , $t < 0.01$
146.1± 2.9	140k	BIGGS	70	CNTR	$< 4.1 \gamma \text{C} \rightarrow \pi^+\pi^-\text{C}$
160 ±10		LANZEROTTI	68	CNTR	$\gamma p$
130 ± 5	4000	ASBURY	67B	CNTR	$\gamma + \text{Pb}$

<sup>1</sup> Using the KUHN 90 parametrization of the pion form factor, neglecting  $\rho$ - $\omega$  interference.

<sup>2</sup> From the parametrization according to SOEDING 66.

<sup>3</sup> From the parametrization according to ROSS 66.

**NEUTRAL ONLY, OTHER REACTIONS**

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>150.9 ± 1.7</b>	<b>OUR AVERAGE</b>	Error includes scale factor of 1.1.		
122 ± 20		BERTIN	97C	OBLX 0.0 $\bar{p}p \rightarrow \pi^+\pi^-\pi^0$
145.7 ± 5.3		WEIDENAUER	93	ASTE $\bar{p}p \rightarrow \pi^+\pi^-\omega$
144.9 ± 3.7		DUBNICKA	89	RVUE $\pi$ form factor
148 ± 6		<sup>1,2</sup> BOHACIK	80	RVUE
152 ± 9		<sup>3</sup> WICKLUND	78	ASPK 3,4,6 $\pi^\pm pN$
154 ± 2	76k	DEUTSCH...	76	HBC 16 $\pi^+p$
157 ± 8	6.8k	<sup>4</sup> RATCLIFF	72	ASPK 15 $\pi^-p$ , $t < 0.3$
143 ± 8	1.7k	REYNOLDS	69	HBC 2.26 $\pi^-p$
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
150.85 ± 0.55 ± 0.67	970k	<sup>5</sup> ABLIKIM	18C	BES3 $\eta'(958) \rightarrow \gamma\pi^+\pi^-$
150.18 ± 0.55 ± 0.65	970k	<sup>6</sup> ABLIKIM	18C	BES3 $\eta'(958) \rightarrow \gamma\pi^+\pi^-$
147.0 ± 2.5	600k	<sup>7</sup> ABELE	99E	CBAR 0.0 $\bar{p}p \rightarrow \pi^+\pi^-\pi^0$
146 ± 3	4.9k	<sup>8</sup> ADAMS	97	E665 470 $\mu p \rightarrow \mu XB$
160.0 <sup>+4.1</sup> <sub>-4.0</sub>		<sup>9</sup> CHABAUD	83	ASPK 17 $\pi^-p$ polarized
155 ± 1		<sup>10</sup> HEYN	81	RVUE $\pi$ form factor
148.0 ± 1.3		<sup>1,2</sup> LANG	79	RVUE
146 ± 14	4.1k	ENGLER	74	DBC 6 $\pi^+n \rightarrow \pi^+\pi^-p$
143 ± 13		<sup>2</sup> ESTABROOKS	74	RVUE 17 $\pi^-p \rightarrow \pi^+\pi^-n$
160 ± 10	32k	<sup>1</sup> PROTOPOP...	73	HBC 7.1 $\pi^+p$ , $t < 0.4$
145 ± 12	2.2k	<sup>3,11</sup> HYAMS	68	OSPK 11.2 $\pi^-p$
163 ± 15	13.3k	<sup>12</sup> PISUT	68	RVUE 1.7–3.2 $\pi^-p$ , $t < 10$

<sup>1</sup> From pole extrapolation.<sup>2</sup> From phase shift analysis of GRAYER 74 data.<sup>3</sup> Width errors enlarged by us to  $4\Gamma/\sqrt{N}$ ; see the note with the  $K^*(892)$  mass.<sup>4</sup> Published values contain misprints. Corrected by private communication RATCLIFF 74.<sup>5</sup> From a fit to  $\pi^+\pi^-$  mass using  $\rho(770)$  (parametrized with the Gounaris-Sakurai approach),  $\omega(782)$ , and box anomaly components.<sup>6</sup> From a fit to  $\pi^+\pi^-$  mass using  $\rho(770)$  (parametrized with the Gounaris-Sakurai approach),  $\omega(782)$ , and  $\rho(1450)$  components.<sup>7</sup> Using relativistic Breit-Wigner and taking into account  $\rho$ - $\omega$  interference.<sup>8</sup> Systematic errors not evaluated.<sup>9</sup> From fit of 3-parameter relativistic Breit-Wigner to helicity-zero part of  $P$ -wave intensity. CHABAUD 83 includes data of GRAYER 74.<sup>10</sup> HEYN 81 includes all spacelike and timelike  $F_\pi$  values until 1978.<sup>11</sup> Of HYAMS 68 six parametrizations this is theoretically soundest. MR<sup>12</sup> Includes MALAMUD 69, ARMENISE 68, BACON 67, HUWE 67, MILLER 67B, ALFF-STEINBERGER 66, HAGOPIAN 66, HAGOPIAN 66B, JACOBS 66B, JAMES 66, WEST 66, GOLDHABER 64, ABOLINS 63. **$\Gamma_{\rho(770)^0} - \Gamma_{\rho(770)^\pm}$** 

VALUE (MeV)	EVTS	DOCUMENT ID	TECN	COMMENT
<b>0.3 ± 1.3</b>	<b>OUR AVERAGE</b>	Error includes scale factor of 1.4.		
-0.2 ± 1.0		<sup>1</sup> SCHAEL	05C	ALEP $\tau^- \rightarrow \pi^-\pi^0\nu_\tau$
3.6 ± 1.8 ± 1.7	1.98M	<sup>2</sup> ALOISIO	03	KLOE 1.02 $e^+e^- \rightarrow \pi^+\pi^-\pi^0$

● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●

4.66±0.85 <sup>3</sup> BARTOS 17A RVUE  $e^+e^- \rightarrow \pi^+\pi^-, \tau^- \rightarrow \pi^-\pi^0\nu_\tau$

<sup>1</sup>From the combined fit of the  $\tau^-$  data from ANDERSON 00A and SCHAEEL 05C and  $e^+e^-$  data from the compilation of BARKOV 85, AKHMETSHIN 04, and ALOISIO 05. Supersedes BARATE 97M.

<sup>2</sup>Assuming  $m_{\rho^+} = m_{\rho^-}$ ,  $\Gamma_{\rho^+} = \Gamma_{\rho^-}$ .

<sup>3</sup>Applies the Unitary & Analytic Model of the pion electromagnetic form factor of DUBNICKA 10 to analyze the data of ACHASOV 06, AKHMETSHIN 07, AUBERT 09AS, AMBROSINO 11A, and FUJIKAWA 08.

$\Gamma_{\rho(770)^+} - \Gamma_{\rho(770)^-}$

VALUE	EVTS	DOCUMENT ID	TECN	COMMENT
<b>1.8±2.0±0.5</b>	1.98M	<sup>1</sup> ALOISIO	03	KLOE 1.02 $e^+e^- \rightarrow \pi^+\pi^-\pi^0$

<sup>1</sup>Without limitations on masses and widths.

### $\rho(770)$ DECAY MODES

Mode	Fraction ( $\Gamma_i/\Gamma$ )	Scale factor/ Confidence level
$\Gamma_1$ $\pi\pi$	~ 100	%
<b><math>\rho(770)^\pm</math> decays</b>		
$\Gamma_2$ $\pi^\pm\pi^0$	~ 100	%
$\Gamma_3$ $\pi^\pm\gamma$	( 4.5 ±0.5 )	$\times 10^{-4}$ S=2.2
$\Gamma_4$ $\pi^\pm\eta$	< 6	$\times 10^{-3}$ CL=84%
$\Gamma_5$ $\pi^\pm\pi^+\pi^-\pi^0$	< 2.0	$\times 10^{-3}$ CL=84%
<b><math>\rho(770)^0</math> decays</b>		
$\Gamma_6$ $\pi^+\pi^-$	~ 100	%
$\Gamma_7$ $\pi^+\pi^-\gamma$	( 9.9 ±1.6 )	$\times 10^{-3}$
$\Gamma_8$ $\pi^0\gamma$	( 4.7 ±0.6 )	$\times 10^{-4}$ S=1.4
$\Gamma_9$ $\eta\gamma$	( 3.00±0.21 )	$\times 10^{-4}$
$\Gamma_{10}$ $\pi^0\pi^0\gamma$	( 4.5 ±0.8 )	$\times 10^{-5}$
$\Gamma_{11}$ $\mu^+\mu^-$	[a] ( 4.55±0.28 )	$\times 10^{-5}$
$\Gamma_{12}$ $e^+e^-$	[a] ( 4.72±0.05 )	$\times 10^{-5}$
$\Gamma_{13}$ $\pi^+\pi^-\pi^0$	( 1.01 <sup>+0.54</sup> <sub>-0.36</sub> ±0.34 )	$\times 10^{-4}$
$\Gamma_{14}$ $\pi^+\pi^-\pi^+\pi^-$	( 1.8 ±0.9 )	$\times 10^{-5}$
$\Gamma_{15}$ $\pi^+\pi^-\pi^0\pi^0$	( 1.6 ±0.8 )	$\times 10^{-5}$
$\Gamma_{16}$ $\pi^0e^+e^-$	< 1.2	$\times 10^{-5}$ CL=90%
$\Gamma_{17}$ $\eta e^+e^-$		

[a] The  $\omega\rho$  interference is then due to  $\omega\rho$  mixing only, and is expected to be small. If  $e\mu$  universality holds,  $\Gamma(\rho^0 \rightarrow \mu^+\mu^-) = \Gamma(\rho^0 \rightarrow e^+e^-) \times 0.99785$ .

### CONSTRAINED FIT INFORMATION

An overall fit to the total width and a partial width uses 10 measurements and one constraint to determine 3 parameters. The overall fit has a  $\chi^2 = 10.7$  for 8 degrees of freedom.

The following *off-diagonal* array elements are the correlation coefficients  $\langle \delta p_i \delta p_j \rangle / (\delta p_i \cdot \delta p_j)$ , in percent, from the fit to parameters  $p_i$ , including the branching fractions,  $x_i \equiv \Gamma_i / \Gamma_{\text{total}}$ . The fit constrains the  $x_i$  whose labels appear in this array to sum to one.

$x_3$	-100	
$\Gamma$	15	-15
	$x_2$	$x_3$

	Mode	Rate (MeV)	Scale factor
$\Gamma_2$	$\pi^\pm \pi^0$	150.2 $\pm$ 2.4	
$\Gamma_3$	$\pi^\pm \gamma$	0.068 $\pm$ 0.007	2.3

### CONSTRAINED FIT INFORMATION

An overall fit to the total width, a partial width, and 7 branching ratios uses 22 measurements and one constraint to determine 9 parameters. The overall fit has a  $\chi^2 = 9.5$  for 14 degrees of freedom.

The following *off-diagonal* array elements are the correlation coefficients  $\langle \delta p_i \delta p_j \rangle / (\delta p_i \cdot \delta p_j)$ , in percent, from the fit to parameters  $p_i$ , including the branching fractions,  $x_i \equiv \Gamma_i / \Gamma_{\text{total}}$ . The fit constrains the  $x_i$  whose labels appear in this array to sum to one.

$x_7$	-100							
$x_8$	-4	0						
$x_9$	-1	0	1					
$x_{10}$	-1	0	0	0				
$x_{11}$	2	-3	0	0	0			
$x_{12}$	0	0	-8	-9	0	0		
$x_{14}$	-1	0	0	0	0	0	0	
$\Gamma$	0	0	4	5	0	0	-54	0
	$x_6$	$x_7$	$x_8$	$x_9$	$x_{10}$	$x_{11}$	$x_{12}$	$x_{14}$

Mode	Rate (MeV)	Scale factor
$\Gamma_6 \quad \pi^+ \pi^-$	147.5 $\pm 0.9$	
$\Gamma_7 \quad \pi^+ \pi^- \gamma$	1.48 $\pm 0.24$	
$\Gamma_8 \quad \pi^0 \gamma$	0.070 $\pm 0.009$	1.4
$\Gamma_9 \quad \eta \gamma$	0.0447 $\pm 0.0032$	
$\Gamma_{10} \quad \pi^0 \pi^0 \gamma$	0.0066 $\pm 0.0012$	
$\Gamma_{11} \quad \mu^+ \mu^-$	[a] 0.0068 $\pm 0.0004$	
$\Gamma_{12} \quad e^+ e^-$	[a] 0.00704 $\pm 0.00006$	
$\Gamma_{14} \quad \pi^+ \pi^- \pi^+ \pi^-$	0.0027 $\pm 0.0014$	

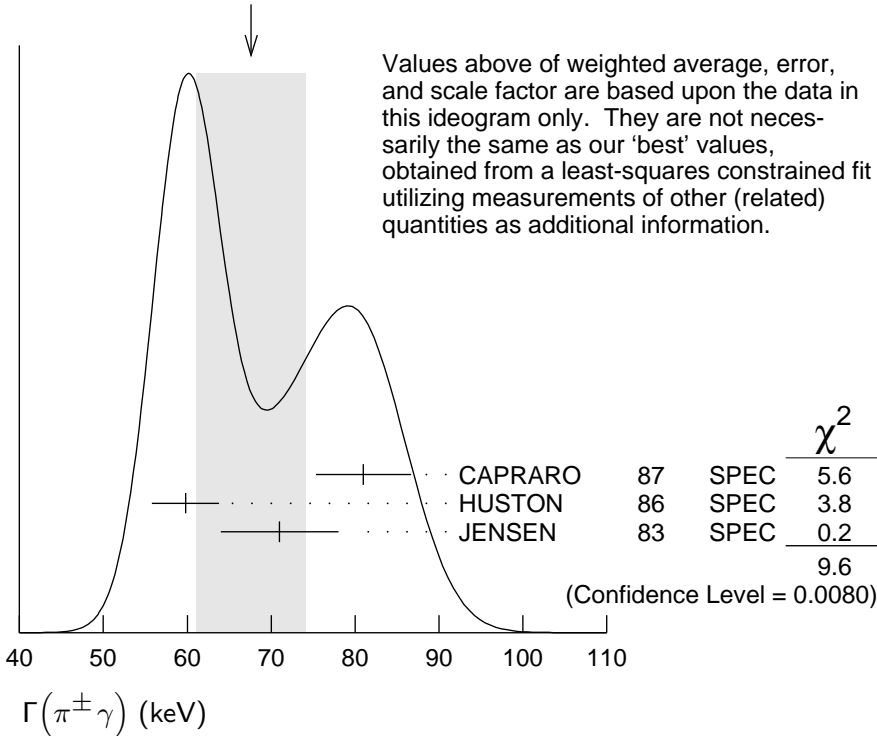
### $\rho(770)$ PARTIAL WIDTHS

$\Gamma(\pi^\pm \gamma)$

$\Gamma_3$

VALUE (keV)	DOCUMENT ID	TECN	CHG	COMMENT
<b>68 <math>\pm 7</math> OUR FIT</b>	Error includes scale factor of 2.3.			
<b>68 <math>\pm 7</math> OUR AVERAGE</b>	Error includes scale factor of 2.2. See the ideogram below.			
81 $\pm 4 \pm 4$	CAPRARO	87	SPEC	- 200 $\pi^- A \rightarrow \pi^- \pi^0 A$
59.8 $\pm 4.0$	HUSTON	86	SPEC	+ 202 $\pi^+ A \rightarrow \pi^+ \pi^0 A$
71 $\pm 7$	JENSEN	83	SPEC	- 156-260 $\pi^- A \rightarrow \pi^- \pi^0 A$

WEIGHTED AVERAGE  
68 $\pm 7$  (Error scaled by 2.2)





$\Gamma(\pi^0\gamma)$  $\Gamma_8$ 

VALUE (keV)	EVTS	DOCUMENT ID	TECN	COMMENT
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• • • We do not use the following data for averages, fits, limits, etc. • • •

$77 \pm 17 \pm 11$	36500	<sup>1</sup> ACHASOV 03	SND	$0.60\text{--}0.97 e^+e^- \rightarrow \pi^0\gamma$
$121 \pm 31$		DOLINSKY 89	ND	$e^+e^- \rightarrow \pi^0\gamma$

<sup>1</sup> Using  $\Gamma_{\text{total}} = 147.9 \pm 1.3$  MeV and  $B(\rho \rightarrow \pi^0\gamma)$  from ACHASOV 03.

 $\Gamma(\eta\gamma)$  $\Gamma_9$ 

VALUE (keV)	DOCUMENT ID	TECN	COMMENT
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• • • We do not use the following data for averages, fits, limits, etc. • • •

$62 \pm 17$	<sup>1</sup> DOLINSKY 89	ND	$e^+e^- \rightarrow \eta\gamma$
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<sup>1</sup> Solution corresponding to constructive  $\omega$ - $\rho$  interference.

 $\Gamma(e^+e^-)$  $\Gamma_{12}$ 

VALUE (keV)	EVTS	DOCUMENT ID	TECN	COMMENT
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**7.04  $\pm$  0.06 OUR FIT**

**7.04  $\pm$  0.06 OUR AVERAGE**

$7.048 \pm 0.057 \pm 0.050$	900k	<sup>1</sup> AKHMETSHIN 07		$e^+e^- \rightarrow \pi^+\pi^-$
$7.06 \pm 0.11 \pm 0.05$	114k	<sup>2,3</sup> AKHMETSHIN 04	CMD2	$e^+e^- \rightarrow \pi^+\pi^-$
$6.77 \pm 0.10 \pm 0.30$		BARKOV 85	OLYA	$e^+e^- \rightarrow \pi^+\pi^-$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$7.12 \pm 0.02 \pm 0.11$	800k	<sup>4</sup> ACHASOV 06	SND	$e^+e^- \rightarrow \pi^+\pi^-$
$6.3 \pm 0.1$		<sup>5</sup> BENAYOUN 98	RVUE	$e^+e^- \rightarrow \pi^+\pi^-, \mu^+\mu^-$

<sup>1</sup> A combined fit of AKHMETSHIN 07, AULCHENKO 06, and AULCHENKO 05.

<sup>2</sup> Using the GOUNARIS 68 parametrization with the complex phase of the  $\rho$ - $\omega$  interference.

<sup>3</sup> From a fit in the energy range 0.61 to 0.96 GeV. Update of AKHMETSHIN 02.

<sup>4</sup> Supersedes ACHASOV 05A.

<sup>5</sup> Using the data of BARKOV 85 in the hidden local symmetry model.

 $\Gamma(\pi^+\pi^-\pi^+\pi^-)$  $\Gamma_{14}$ 

VALUE (keV)	EVTS	DOCUMENT ID	TECN	COMMENT
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• • • We do not use the following data for averages, fits, limits, etc. • • •

$2.8 \pm 1.4 \pm 0.5$	153	AKHMETSHIN 00	CMD2	$0.6\text{--}0.97 e^+e^- \rightarrow \pi^+\pi^-\pi^+\pi^-$
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 $\rho(770) \Gamma(e^+e^-)\Gamma(i)/\Gamma^2(\text{total})$  $\Gamma(e^+e^-)/\Gamma_{\text{total}} \times \Gamma(\pi^+\pi^-)/\Gamma_{\text{total}}$  $\Gamma_{12}/\Gamma \times \Gamma_6/\Gamma$ 

VALUE (units $10^{-5}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
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**4.876  $\pm$  0.023  $\pm$  0.064** 800k <sup>1,2</sup>ACHASOV 06 SND  $e^+e^- \rightarrow \pi^+\pi^-$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$4.72 \pm 0.02$		<sup>3</sup> BENAYOUN 10	RVUE	$0.4\text{--}1.05 e^+e^-$
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<sup>1</sup> Supersedes ACHASOV 05A.

<sup>2</sup> A fit of the SND data from 400 to 1000 MeV using parameters of the  $\rho(1450)$  and  $\rho(1700)$  from a fit of the data of BARKOV 85, BISELLO 89 and ANDERSON 00A.

<sup>3</sup> A simultaneous fit of  $e^+e^- \rightarrow \pi^+\pi^-, \pi^+\pi^-\pi^0, \pi^0\gamma, \eta\gamma$  data.

$$\Gamma(e^+e^-)/\Gamma_{\text{total}} \times \Gamma(\eta\gamma)/\Gamma_{\text{total}} \qquad \Gamma_{12}/\Gamma \times \Gamma_9/\Gamma$$

VALUE (units $10^{-8}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
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**1.42±0.10 OUR FIT****1.45±0.12 OUR AVERAGE**

1.32±0.14±0.08	33k	<sup>1</sup> ACHASOV	07B SND	0.6–1.38 $e^+e^- \rightarrow \eta\gamma$
1.50±0.65±0.09	17.4k	<sup>2</sup> AKHMETSHIN	05 CMD2	0.60–1.38 $e^+e^- \rightarrow \eta\gamma$
1.61±0.20±0.11	23k	<sup>3,4</sup> AKHMETSHIN	01B CMD2	$e^+e^- \rightarrow \eta\gamma$
1.85±0.49		<sup>5</sup> DOLINSKY	89 ND	$e^+e^- \rightarrow \eta\gamma$

● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●

1.05±0.02 <sup>6</sup>BENAYOUN 10 RVUE 0.4–1.05  $e^+e^-$ 

<sup>1</sup> From a combined fit of  $\sigma(e^+e^- \rightarrow \eta\gamma)$  with  $\eta \rightarrow 3\pi^0$  and  $\eta \rightarrow \pi^+\pi^-\pi^0$ , and fixing  $B(\eta \rightarrow 3\pi^0) / B(\eta \rightarrow \pi^+\pi^-\pi^0) = 1.44 \pm 0.04$ . Recalculated by us from the cross section at the peak. Supersedes ACHASOV 00D and ACHASOV 06A.

<sup>2</sup> From the  $\eta \rightarrow 2\gamma$  decay and using  $B(\eta \rightarrow \gamma\gamma) = 39.43 \pm 0.26\%$ .

<sup>3</sup> From the  $\eta \rightarrow 3\pi^0$  decay and using  $B(\eta \rightarrow 3\pi^0) = (32.24 \pm 0.29) \times 10^{-2}$ .

<sup>4</sup> The combined fit from 600 to 1380 MeV taking into account  $\rho(770)$ ,  $\omega(782)$ ,  $\phi(1020)$ , and  $\rho(1450)$  (mass and width fixed at 1450 MeV and 310 MeV respectively).

<sup>5</sup> Recalculated by us from the cross section in the peak.

<sup>6</sup> A simultaneous fit of  $e^+e^- \rightarrow \pi^+\pi^-, \pi^+\pi^-\pi^0, \pi^0\gamma, \eta\gamma$  data.

$$\Gamma(e^+e^-)/\Gamma_{\text{total}} \times \Gamma(\pi^0\gamma)/\Gamma_{\text{total}} \qquad \Gamma_{12}/\Gamma \times \Gamma_8/\Gamma$$

VALUE (units $10^{-8}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
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**2.22 ±0.29 OUR FIT** Error includes scale factor of 1.4.**2.22 ±0.26 OUR AVERAGE** Error includes scale factor of 1.3. See the ideogram below.

1.98 ±0.22 ±0.10		<sup>1</sup> ACHASOV	16A SND	0.60–1.38 $e^+e^- \rightarrow \pi^0\gamma$
2.90 $^{+0.60}_{-0.55}$ ±0.18	18k	AKHMETSHIN	05 CMD2	0.60–1.38 $e^+e^- \rightarrow \pi^0\gamma$
2.37 ±0.53 ±0.33	36k	<sup>2</sup> ACHASOV	03 SND	0.60–0.97 $e^+e^- \rightarrow \pi^0\gamma$
3.61 ±0.74 ±0.49	10k	<sup>3</sup> DOLINSKY	89 ND	$e^+e^- \rightarrow \pi^0\gamma$

● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●

1.875±0.026 <sup>4</sup>BENAYOUN 10 RVUE 0.4–1.05  $e^+e^-$ 

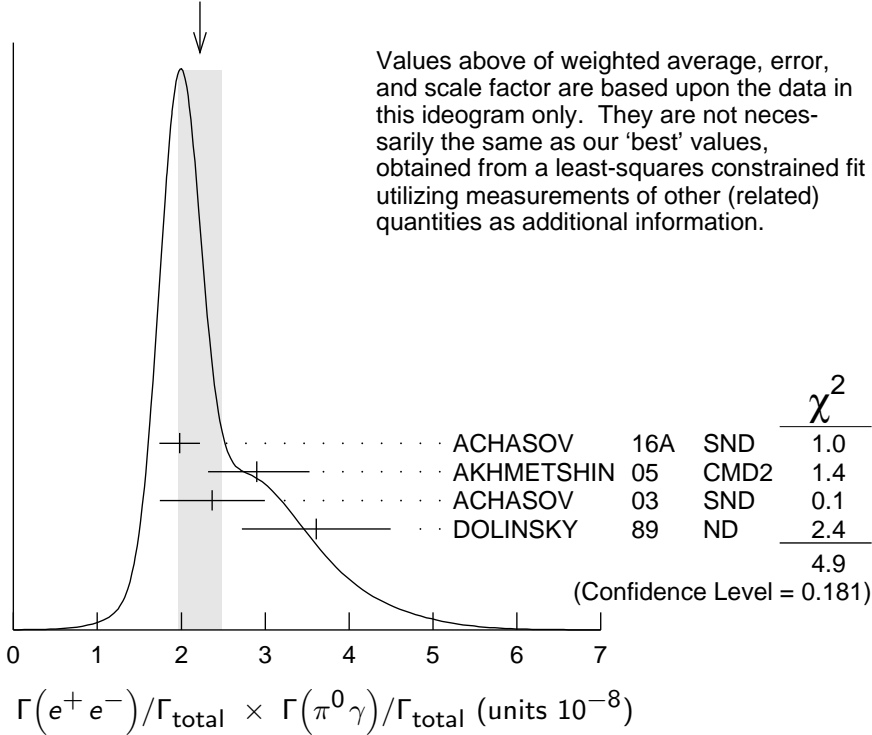
<sup>1</sup> From the VMD model with the  $\rho(770)$ ,  $\omega(782)$ ,  $\phi(1020)$  resonances, and an additional resonance describing the total contribution of the  $\rho(1450)$  and  $\omega(1420)$  states. Supersedes ACHASOV 03.

<sup>2</sup> Using  $\sigma_{\phi \rightarrow \pi^0\gamma}$  from ACHASOV 00 and  $m_\rho = 775.97$  MeV in the model with the energy-independent phase of  $\rho$ - $\omega$  interference equal to  $(-10.2 \pm 7.0)^\circ$ .

<sup>3</sup> Recalculated by us from the cross section in the peak.

<sup>4</sup> A simultaneous fit of  $e^+e^- \rightarrow \pi^+\pi^-, \pi^+\pi^-\pi^0, \pi^0\gamma, \eta\gamma$  data.

WEIGHTED AVERAGE  
2.22±0.26 (Error scaled by 1.3)



**$\Gamma(e^+e^-)/\Gamma_{\text{total}} \times \Gamma(\pi^+\pi^-\pi^0)/\Gamma_{\text{total}}$   $\Gamma_{12}/\Gamma \times \Gamma_{13}/\Gamma$**

VALUE (units $10^{-9}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
0.903±0.076		<sup>1</sup> BENAYOUN 10	RVUE	0.4–1.05 $e^+e^-$
4.58 $\begin{smallmatrix} +2.46 \\ -1.64 \end{smallmatrix}$ ±1.56	1.2M	<sup>2</sup> ACHASOV 03D	RVUE	0.44–2.00 $e^+e^- \rightarrow \pi^+\pi^-\pi^0$

<sup>1</sup> A simultaneous fit of  $e^+e^- \rightarrow \pi^+\pi^-, \pi^+\pi^-\pi^0, \pi^0\gamma, \eta\gamma$  data.  
<sup>2</sup> Statistical significance is less than  $3\sigma$ .

**$\rho(770)$  BRANCHING RATIOS**

**$\Gamma(\pi^\pm\eta)/\Gamma(\pi\pi)$   $\Gamma_4/\Gamma_1$**

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	CHG	COMMENT
<b>&lt;60</b>	84	FERBEL 66	HBC	±	$\pi^\pm p$ above 2.5

**$\Gamma(\pi^\pm\pi^+\pi^-\pi^0)/\Gamma(\pi\pi)$   $\Gamma_5/\Gamma_1$**

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	CHG	COMMENT
<b>&lt;20</b>	84	FERBEL 66	HBC	±	$\pi^\pm p$ above 2.5

● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●

35±40		JAMES 66	HBC	+	2.1 $\pi^+ p$
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$\Gamma(\pi^+\pi^-\gamma)/\Gamma_{\text{total}}$   $\Gamma_7/\Gamma$ 

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
<b>0.0099±0.0016 OUR FIT</b>				
<b>0.0099±0.0016</b>		<sup>1</sup> DOLINSKY 91	ND	$e^+e^- \rightarrow \pi^+\pi^-\gamma$
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.0111±0.0014		<sup>2</sup> VASSERMAN 88	ND	$e^+e^- \rightarrow \pi^+\pi^-\gamma$
<0.005	90	<sup>3</sup> VASSERMAN 88	ND	$e^+e^- \rightarrow \pi^+\pi^-\gamma$

<sup>1</sup> Bremsstrahlung from a decay pion and for photon energy above 50 MeV.<sup>2</sup> Superseded by DOLINSKY 91.<sup>3</sup> Structure radiation due to quark rearrangement in the decay. $\Gamma(\pi^0\gamma)/\Gamma_{\text{total}}$   $\Gamma_8/\Gamma$ 

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	COMMENT
• • • We do not use the following data for averages, fits, limits, etc. • • •				
4.20±0.52		<sup>1</sup> ACHASOV 16A	SND	0.60–1.38 $e^+e^- \rightarrow \pi^0\gamma$
$6.21^{+1.28}_{-1.18} \pm 0.39$	18k	<sup>2,3</sup> AKHMETSHIN 05	CMD2	0.60–1.38 $e^+e^- \rightarrow \pi^0\gamma$
$5.22 \pm 1.17 \pm 0.75$	36k	<sup>3,4</sup> ACHASOV 03	SND	0.60–0.97 $e^+e^- \rightarrow \pi^0\gamma$
$6.8 \pm 1.7$		<sup>5</sup> BENAYOUN 96	RVUE	0.54–1.04 $e^+e^- \rightarrow \pi^0\gamma$
$7.9 \pm 2.0$		<sup>3</sup> DOLINSKY 89	ND	$e^+e^- \rightarrow \pi^0\gamma$

<sup>1</sup> Using  $B(\rho \rightarrow e^+e^-)$  from PDG 15. Supersedes ACHASOV 03.<sup>2</sup> Using  $B(\rho \rightarrow e^+e^-) = (4.67 \pm 0.09) \times 10^{-5}$ .<sup>3</sup> Not independent of the corresponding  $\Gamma(e^+e^-) \times \Gamma(\pi^0\gamma)/\Gamma_{\text{total}}^2$ .<sup>4</sup> Using  $B(\rho \rightarrow e^+e^-) = (4.54 \pm 0.10) \times 10^{-5}$ .<sup>5</sup> Reanalysis of DRUZHININ 84, DOLINSKY 89, and DOLINSKY 91 taking into account a triangle anomaly contribution. $\Gamma(\eta\gamma)/\Gamma_{\text{total}}$   $\Gamma_9/\Gamma$ 

VALUE (units $10^{-4}$ )	EVTS	DOCUMENT ID	TECN	CHG	COMMENT
<b>3.00±0.21 OUR FIT</b>					
<b>2.90±0.32 OUR AVERAGE</b>					
$2.79 \pm 0.34 \pm 0.03$	33k	<sup>1</sup> ACHASOV 07B	SND		0.6–1.38 $e^+e^- \rightarrow \eta\gamma$
$3.6 \pm 0.9$		<sup>2</sup> ANDREWS 77	CNTR	0	6.7–10 $\gamma\text{Cu}$
• • • We do not use the following data for averages, fits, limits, etc. • • •					
$3.21 \pm 1.39 \pm 0.20$	17.4k	<sup>3,4</sup> AKHMETSHIN 05	CMD2		0.60–1.38 $e^+e^- \rightarrow \eta\gamma$
$3.39 \pm 0.42 \pm 0.23$		<sup>2,5,6</sup> AKHMETSHIN 01B	CMD2		$e^+e^- \rightarrow \eta\gamma$
$1.9^{+0.6}_{-0.8}$		<sup>7</sup> BENAYOUN 96	RVUE		0.54–1.04 $e^+e^- \rightarrow \eta\gamma$
$4.0 \pm 1.1$		<sup>2,4</sup> DOLINSKY 89	ND		$e^+e^- \rightarrow \eta\gamma$

<sup>1</sup> ACHASOV 07B reports  $[\Gamma(\rho(770) \rightarrow \eta\gamma)/\Gamma_{\text{total}}] \times [B(\rho(770) \rightarrow e^+e^-)] = (1.32 \pm 0.14 \pm 0.08) \times 10^{-8}$  which we divide by our best value  $B(\rho(770) \rightarrow e^+e^-) = (4.72 \pm 0.05) \times 10^{-5}$ . Our first error is their experiment's error and our second error is the systematic error from using our best value. Supersedes ACHASOV 00D and ACHASOV 06A.<sup>2</sup> Solution corresponding to constructive  $\omega$ - $\rho$  interference.<sup>3</sup> Using  $B(\rho \rightarrow e^+e^-) = (4.67 \pm 0.09) \times 10^{-5}$  and  $B(\eta \rightarrow \gamma\gamma) = 39.43 \pm 0.26\%$ .<sup>4</sup> Not independent of the corresponding  $\Gamma(e^+e^-) \times \Gamma(\eta\gamma)/\Gamma_{\text{total}}^2$ .<sup>5</sup> The combined fit from 600 to 1380 MeV taking into account  $\rho(770)$ ,  $\omega(782)$ ,  $\phi(1020)$ , and  $\rho(1450)$  (mass and width fixed at 1450 MeV and 310 MeV respectively).

<sup>6</sup> Using  $B(\rho \rightarrow e^+ e^-) = (4.75 \pm 0.10) \times 10^{-5}$  from AKHMETSHIN 02 and  $B(\eta \rightarrow 3\pi^0) = (32.24 \pm 0.29) \times 10^{-2}$ .

<sup>7</sup> Reanalysis of DRUZHININ 84, DOLINSKY 89, and DOLINSKY 91 taking into account a triangle anomaly contribution. Constructive  $\rho$ - $\omega$  interference solution.

### $\Gamma(\pi^0 \pi^0 \gamma) / \Gamma_{\text{total}}$ $\Gamma_{10} / \Gamma$

<u>VALUE (units <math>10^{-5}</math>)</u>	<u>EPTS</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
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**4.5 ± 0.8 OUR FIT**

**4.5<sup>+0.9</sup><sub>-0.8</sub> OUR AVERAGE**

5.2 <sup>+1.5</sup> <sub>-1.3</sub> ± 0.6	190	<sup>1</sup> AKHMETSHIN 04B	CMD2	0.6–0.97 $e^+ e^- \rightarrow \pi^0 \pi^0 \gamma$
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4.1 <sup>+1.0</sup> <sub>-0.9</sub> ± 0.3	295	<sup>2</sup> ACHASOV	02F SND	0.36–0.97 $e^+ e^- \rightarrow \pi^0 \pi^0 \gamma$
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• • • We do not use the following data for averages, fits, limits, etc. • • •

4.8 <sup>+3.4</sup> <sub>-1.8</sub> ± 0.5	63	<sup>3</sup> ACHASOV	00G SND	$e^+ e^- \rightarrow \pi^0 \pi^0 \gamma$
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<sup>1</sup> This branching ratio includes the conventional VMD mechanism  $\rho \rightarrow \omega \pi^0$ ,  $\omega \rightarrow \pi^0 \gamma$ , and the new decay mode  $\rho \rightarrow f_0(500) \gamma$ ,  $f_0(500) \rightarrow \pi^0 \pi^0$  with a branching ratio  $(2.0^{+1.1}_{-0.9} \pm 0.3) \times 10^{-5}$  differing from zero by 2.0 standard deviations.

<sup>2</sup> This branching ratio includes the conventional VMD mechanism  $\rho \rightarrow \omega \pi^0$ ,  $\omega \rightarrow \pi^0 \gamma$  and the new decay mode  $\rho \rightarrow f_0(500) \gamma$ ,  $f_0(500) \rightarrow \pi^0 \pi^0$  with a branching ratio  $(1.9^{+0.9}_{-0.8} \pm 0.4) \times 10^{-5}$  differing from zero by 2.4 standard deviations. Supersedes ACHASOV 00G.

<sup>3</sup> Superseded by ACHASOV 02F.

### $\Gamma(\mu^+ \mu^-) / \Gamma(\pi^+ \pi^-)$ $\Gamma_{11} / \Gamma_6$

<u>VALUE (units <math>10^{-5}</math>)</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
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**4.60 ± 0.28 OUR FIT**

<b>4.6 ± 0.2 ± 0.2</b>	ANTIPOV	89	SIGM $\pi^- \text{Cu} \rightarrow \mu^+ \mu^- \pi^- \text{Cu}$
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• • • We do not use the following data for averages, fits, limits, etc. • • •

8.2 <sup>+1.6</sup> <sub>-3.6</sub>	<sup>1</sup> ROTHWELL	69	CNTR Photoproduction
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5.6 ± 1.5	<sup>2</sup> WEHMANN	69	OSPK 12 $\pi^- \text{C, Fe}$
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9.7 <sup>+3.1</sup> <sub>-3.3</sub>	<sup>3,4</sup> HYAMS	67	OSPK 11 $\pi^- \text{Li, H}$
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<sup>1</sup> Possibly large  $\rho$ - $\omega$  interference leads us to increase the minus error.

<sup>2</sup> Result contains  $11 \pm 11\%$  correction using SU(3) for central value. The error on the correction takes account of possible  $\rho$ - $\omega$  interference and the upper limit agrees with the upper limit of  $\omega \rightarrow \mu^+ \mu^-$  from this experiment.

<sup>3</sup> But he even enlarges his error to take residual  $\omega$  contamination into account. Since his value is high, seems the other experiments also can't have too many  $\omega$ 's. But maybe Hyams has additional  $\mu$ 's from  $\rho \rightarrow \pi \pi$ , decaying  $\pi$ 's.

<sup>4</sup> HYAMS 67's mass resolution is 20 MeV. The  $\omega$  region was excluded.

### $\Gamma(e^+ e^-) / \Gamma(\pi \pi)$ $\Gamma_{12} / \Gamma_1$

<u>VALUE (units <math>10^{-4}</math>)</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
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• • • We do not use the following data for averages, fits, limits, etc. • • •

0.40 ± 0.05	<sup>1,2</sup> BENAKSAS	72	OSPK $e^+ e^- \rightarrow \pi^+ \pi^-$
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<sup>1</sup> The  $\rho'$  contribution is not taken into account.

<sup>2</sup> Barkov excludes Auslender and Benaksas for large statistical and systematic errors.

$\Gamma(\pi^+\pi^-\pi^0)/\Gamma_{\text{total}}$   $\Gamma_{13}/\Gamma$

VALUE (units $10^{-4}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
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• • • We do not use the following data for averages, fits, limits, etc. • • •

$1.01^{+0.54}_{-0.36} \pm 0.34$		1.2M	<sup>1</sup> ACHASOV	03D RVUE	$0.44-2.00 e^+e^- \rightarrow \pi^+\pi^-\pi^0$
<1.2	90		VASSERMAN	88B ND	$e^+e^- \rightarrow \pi^+\pi^-\pi^0$

<sup>1</sup>Statistical significance is less than  $3\sigma$ .

$\Gamma(\pi^+\pi^-\pi^0)/\Gamma(\pi\pi)$   $\Gamma_{13}/\Gamma_1$

VALUE	CL%	DOCUMENT ID	TECN	CHG	COMMENT
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• • • We do not use the following data for averages, fits, limits, etc. • • •

$\sim 0.01$		BRAMON	86 RVUE	0	$J/\psi \rightarrow \omega\pi^0$
<0.01	84	<sup>1</sup> ABRAMS	71 HBC	0	$3.7 \pi^+p$

<sup>1</sup>Model dependent, assumes  $l = 1, 2, \text{ or } 3$  for the  $3\pi$  system.

$\Gamma(\pi^+\pi^-\pi^+\pi^-)/\Gamma_{\text{total}}$   $\Gamma_{14}/\Gamma$

VALUE (units $10^{-5}$ )	CL%	EVTS	DOCUMENT ID	TECN	COMMENT
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**1.8±0.9 OUR FIT**

<b>1.8±0.9±0.3</b>		153	AKHMETSHIN 00	CMD2	$0.6-0.97 e^+e^- \rightarrow \pi^+\pi^-\pi^+\pi^-$
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• • • We do not use the following data for averages, fits, limits, etc. • • •

<20	90		KURDADZE	88 OLYA	$e^+e^- \rightarrow \pi^+\pi^-\pi^+\pi^-$
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$\Gamma(\pi^+\pi^-\pi^+\pi^-)/\Gamma(\pi\pi)$   $\Gamma_{14}/\Gamma_1$

VALUE (units $10^{-4}$ )	CL%	DOCUMENT ID	TECN	CHG	COMMENT
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• • • We do not use the following data for averages, fits, limits, etc. • • •

<15	90	ERBE	69 HBC	0	$2.5-5.8 \gamma p$
<20		CHUNG	68 HBC	0	$3.2, 4.2 \pi^- p$
<20	90	HUSON	68 HLBC	0	$16.0 \pi^- p$
<80		JAMES	66 HBC	0	$2.1 \pi^+ p$

$\Gamma(\pi^+\pi^-\pi^0\pi^0)/\Gamma_{\text{total}}$   $\Gamma_{15}/\Gamma$

VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT
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<b>1.60±0.74±0.18</b>		<sup>1</sup> ACHASOV	09A SND	$e^+e^- \rightarrow \pi^+\pi^-\pi^0\pi^0$
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• • • We do not use the following data for averages, fits, limits, etc. • • •

< 4	90	AULCHENKO	87C ND	$e^+e^- \rightarrow \pi^+\pi^-\pi^0\pi^0$
<20	90	KURDADZE	86 OLYA	$e^+e^- \rightarrow \pi^+\pi^-\pi^0\pi^0$

<sup>1</sup>Assuming no interference between the  $\rho$  and  $\omega$  contributions.

$\Gamma(\pi^0 e^+ e^-)/\Gamma_{\text{total}}$   $\Gamma_{16}/\Gamma$

VALUE (units $10^{-5}$ )	CL%	DOCUMENT ID	TECN	COMMENT
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<1.2	90	ACHASOV	08 SND	$0.36-0.97 e^+e^- \rightarrow \pi^0 e^+e^-$
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• • • We do not use the following data for averages, fits, limits, etc. • • •

<1.6		AKHMETSHIN 05A	CMD2	$0.72-0.84 e^+e^-$
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$\Gamma(\eta e^+ e^-)/\Gamma_{\text{total}}$  $\Gamma_{17}/\Gamma$ VALUE (units  $10^{-5}$ )

DOCUMENT ID

TECN

COMMENT

● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●

&lt;0.7

AKHMETSHIN 05A CMD2 0.72-0.84  $e^+ e^-$  **$\rho(770)$  REFERENCES**

ABLIKIM	18C	PRL 120 242003	M. Ablikim <i>et al.</i>	(BESIII Collab.)
ABLIKIM	17	PRL 118 012001	M. Ablikim <i>et al.</i>	(BESIII Collab.)
BARTOS	17	PR D96 113004	E. Bartos <i>et al.</i>	
BARTOS	17A	IJMP A32 1750154	E. Bartos <i>et al.</i>	
ABLIKIM	16C	PL B753 629	M. Ablikim <i>et al.</i>	(BESIII Collab.)
ACHASOV	16A	PR D93 092001	M.N. Achasov <i>et al.</i>	(SND Collab.)
PDG	15	RPP 2015 at pdg.lbl.gov		(PDG Collab.)
ABRAMOWICZ	12	EPJ C72 1869	H. Abramowicz <i>et al.</i>	(ZEUS Collab.)
LEES	12G	PR D86 032013	J.P. Lees <i>et al.</i>	(BABAR Collab.)
AMBROSINO	11A	PL B700 102	F. Ambrosino <i>et al.</i>	(KLOE Collab.)
BENAYOUN	10	EPJ C65 211	M. Benayoun <i>et al.</i>	
DUBNICKA	10	APS 60 1	S. Dubnicka, A.Z. Dubnickova	
ACHASOV	09A	JETP 109 379	M.N. Achasov <i>et al.</i>	(SND Collab.)
		Translated from ZETF 136 442.		
AUBERT	09AS	PRL 103 231801	B. Aubert <i>et al.</i>	(BABAR Collab.)
ACHASOV	08	JETP 107 61	M.N. Achasov <i>et al.</i>	(SND Collab.)
		Translated from ZETF 134 80.		
FUJIKAWA	08	PR D78 072006	M. Fujikawa <i>et al.</i>	(BELLE Collab.)
ACHASOV	07B	PR D76 077101	M.N. Achasov <i>et al.</i>	(SND Collab.)
AKHMETSHIN	07	PL B648 28	R.R. Akhmetshin <i>et al.</i>	(Novosibirsk CMD-2 Collab.)
ACHASOV	06	JETP 103 380	M.N. Achasov <i>et al.</i>	(Novosibirsk SND Collab.)
		Translated from ZETF 130 437.		
ACHASOV	06A	PR D74 014016	M.N. Achasov <i>et al.</i>	(SND Collab.)
AULCHENKO	06	JETPL 84 413	V.M. Aulchenko <i>et al.</i>	(Novosibirsk CMD-2 Collab.)
		Translated from ZETFP 84 491.		
ACHASOV	05A	JETP 101 1053	M.N. Achasov <i>et al.</i>	(Novosibirsk SND Collab.)
		Translated from ZETF 128 1201.		
AKHMETSHIN	05	PL B605 26	R.R. Akhmetshin <i>et al.</i>	(Novosibirsk CMD-2 Collab.)
AKHMETSHIN	05A	PL B613 29	R.R. Akhmetshin <i>et al.</i>	(Novosibirsk CMD-2 Collab.)
ALOISIO	05	PL B606 12	A. Aloisio <i>et al.</i>	(KLOE Collab.)
AULCHENKO	05	JETPL 82 743	V.M. Aulchenko <i>et al.</i>	(Novosibirsk CMD-2 Collab.)
		Translated from ZETFP 82 841.		
SCHAEEL	05C	PRPL 421 191	S. Schaeel <i>et al.</i>	(ALEPH Collab.)
AKHMETSHIN	04	PL B578 285	R.R. Akhmetshin <i>et al.</i>	(Novosibirsk CMD-2 Collab.)
AKHMETSHIN	04B	PL B580 119	R.R. Akhmetshin <i>et al.</i>	(Novosibirsk CMD-2 Collab.)
ACHASOV	03	PL B559 171	M.N. Achasov <i>et al.</i>	(Novosibirsk SND Collab.)
ACHASOV	03D	PR D68 052006	M.N. Achasov <i>et al.</i>	(Novosibirsk SND Collab.)
ALOISIO	03	PL B561 55	A. Aloisio <i>et al.</i>	(KLOE Collab.)
SANZ-CILLERO	03	EPJ C27 587	J.J. Sanz-Cillero, A. Pich	
ACHASOV	02	PR D65 032002	M.N. Achasov <i>et al.</i>	(Novosibirsk SND Collab.)
ACHASOV	02F	PL B537 201	M.N. Achasov <i>et al.</i>	(Novosibirsk SND Collab.)
AKHMETSHIN	02	PL B527 161	R.R. Akhmetshin <i>et al.</i>	(Novosibirsk CMD-2 Collab.)
AKHMETSHIN	01B	PL B509 217	R.R. Akhmetshin <i>et al.</i>	(Novosibirsk CMD-2 Collab.)
COLANGELO	01	NP B603 125	G. Colangelo, J. Gasser, H. Leytwyler	
PICH	01	PR D63 093005	A. Pich, J. Portoles	
ACHASOV	00	EPJ C12 25	M.N. Achasov <i>et al.</i>	(Novosibirsk SND Collab.)
ACHASOV	00D	JETPL 72 282	M.N. Achasov <i>et al.</i>	(Novosibirsk SND Collab.)
		Translated from ZETFP 72 411.		
ACHASOV	00G	JETPL 71 355	M.N. Achasov <i>et al.</i>	(Novosibirsk SND Collab.)
		Translated from ZETFP 71 519.		
AKHMETSHIN	00	PL B475 190	R.R. Akhmetshin <i>et al.</i>	(Novosibirsk CMD-2 Collab.)
ANDERSON	00A	PR D61 112002	S. Anderson <i>et al.</i>	(CLEO Collab.)
ABELE	99E	PL B469 270	A. Abele <i>et al.</i>	(Crystal Barrel Collab.)
BENAYOUN	98	EPJ C2 269	M. Benayoun <i>et al.</i>	(IPNP, NOVO, ADLD+)
BREITWEG	98B	EPJ C2 247	J. Breitweg <i>et al.</i>	(ZEUS Collab.)
GARDNER	98	PR D57 2716	S. Gardner, H.B. O'Connell	
	Also	PR D62 019903 (errat.)	S. Gardner, H.B. O'Connell	
ABELE	97	PL B391 191	A. Abele <i>et al.</i>	(Crystal Barrel Collab.)
ADAMS	97	ZPHY C74 237	M.R. Adams <i>et al.</i>	(E665 Collab.)
BARATE	97M	ZPHY C76 15	R. Barate <i>et al.</i>	(ALEPH Collab.)
BERTIN	97C	PL B408 476	A. Bertin <i>et al.</i>	(OBELIX Collab.)
BOGOLYUB...	97	PAN 60 46	M.Y. Bogolyubsky <i>et al.</i>	(MOSU, SERP)
		Translated from YAF 60 53.		

O'CONNELL	97	NP A623 559	H.B. O'Connell <i>et al.</i>	(ADLD)
BENAYOUN	96	ZPHY C72 221	M. Benayoun <i>et al.</i>	(IPNP, NOVO)
BERNICA	94	PR D50 4454	A. Bernicha, G. Lopez Castro, J. Pestieau	(LOUV+)
WEIDENAUER	93	ZPHY C59 387	P. Weidenauer <i>et al.</i>	(ASTERIX Collab.)
AGUILAR-...	91	ZPHY C50 405	M. Aguilar-Benitez <i>et al.</i>	(LEBC-EHS Collab.)
DOLINSKY	91	PRPL 202 99	S.I. Dolinsky <i>et al.</i>	(NOVO)
KUHN	90	ZPHY C48 445	J.H. Kuhn <i>et al.</i>	(MPIM)
ANTIPOV	89	ZPHY C42 185	Y.M. Antipov <i>et al.</i>	(SERP, JINR, BGNA+)
BISELLO	89	PL B220 321	D. Bisello <i>et al.</i>	(DM2 Collab.)
DOLINSKY	89	ZPHY C42 511	S.I. Dolinsky <i>et al.</i>	(NOVO)
DUBNICKA	89	JP G15 1349	S. Dubnicka <i>et al.</i>	(JINR, SLOV)
GESHKEN...	89	ZPHY C45 351	B.V. Geshkenbein	(ITEP)
KURDADZE	88	JETPL 47 512	L.M. Kurdadze <i>et al.</i>	(NOVO)
		Translated from ZETFP 47 432.		
VASSERMAN	88	SJNP 47 1035	I.B. Vasserman <i>et al.</i>	(NOVO)
		Translated from YAF 47 1635.		
VASSERMAN	88B	SJNP 48 480	I.B. Vasserman <i>et al.</i>	(NOVO)
		Translated from YAF 48 753.		
AULCHENKO	87C	IYF 87-90 Preprint	V.M. Aulchenko <i>et al.</i>	(NOVO)
CAPRARO	87	NP B288 659	L. Capraro <i>et al.</i>	(CLER, FRAS, MILA+)
BRAMON	86	PL B173 97	A. Bramon, J. Casulleras	(BARC)
HUSTON	86	PR D33 3199	J. Huston <i>et al.</i>	(ROCH, FNAL, MINN)
KURDADZE	86	JETPL 43 643	L.M. Kurdadze <i>et al.</i>	(NOVO)
		Translated from ZETFP 43 497.		
BARKOV	85	NP B256 365	L.M. Barkov <i>et al.</i>	(NOVO)
DRUZHININ	84	PL 144B 136	V.P. Druzhinin <i>et al.</i>	(NOVO)
CHABAUD	83	NP B223 1	V. Chabaud <i>et al.</i>	(CERN, CRAC, MPIM)
JENSEN	83	PR D27 26	T. Jensen <i>et al.</i>	(ROCH, FNAL, MINN)
HEYN	81	ZPHY C7 169	M.F. Heyn, C.B. Lang	(GRAZ)
BOHACIK	80	PR D21 1342	J. Bohacik, H. Kuhnelt	(SLOV, WIEN)
LANG	79	PR D19 956	C.B. Lang, A. Mas-Parareda	(GRAZ)
BARTALUCCI	78	NC 44A 587	S. Bartalucci <i>et al.</i>	(DESY, FRAS)
WICKLUND	78	PR D17 1197	A.B. Wicklund <i>et al.</i>	(ANL)
ANDREWS	77	PRL 38 198	D.E. Andrews <i>et al.</i>	(ROCH)
DEUTSCH...	76	NP B103 426	M. Deutschmann <i>et al.</i>	(AACH3, BERL, BONN+)
ENGLER	74	PR D10 2070	A. Engler <i>et al.</i>	(CMU, CASE)
ESTABROOKS	74	NP B79 301	P.G. Estabrooks, A.D. Martin	(DURH)
GRAYER	74	NP B75 189	G. Grayer <i>et al.</i>	(CERN, MPIM)
RATCLIFF	74	Private Comm.		
BYERLY	73	PR D7 637	W.L. Byerly <i>et al.</i>	(MICH)
GLADDING	73	PR D8 3721	G.E. Gladding <i>et al.</i>	(HARV)
HYAMS	73	NP B64 134	B.D. Hyams <i>et al.</i>	(CERN, MPIM)
PROTOPOP...	73	PR D7 1279	S.D. Protopopescu <i>et al.</i>	(LBL)
BALLAM	72	PR D5 545	J. Ballam <i>et al.</i>	(SLAC, LBL, TUFTS)
BENAKSAS	72	PL 39B 289	D. Benaksas <i>et al.</i>	(ORSAY)
JACOBS	72	PR D6 1291	L.D. Jacobs	(SACL)
RATCLIFF	72	PL 38B 345	B.N. Ratcliff <i>et al.</i>	(SLAC)
ABRAMS	71	PR D4 653	G.S. Abrams <i>et al.</i>	(LBL)
ALVENSLEB...	70	PRL 24 786	H. Alvensleben <i>et al.</i>	(DESY)
BIGGS	70	PRL 24 1197	P.J. Biggs <i>et al.</i>	(DARE)
ERBE	69	PR 188 2060	R. Erbe <i>et al.</i>	(German Bubble Chamber Collab.)
MALAMUD	69	Argonne Conf. 93	E.I. Malamud, P.E. Schlein	(UCLA)
REYNOLDS	69	PR 184 1424	B.G. Reynolds <i>et al.</i>	(FSU)
ROTHWELL	69	PRL 23 1521	P.L. Rothwell <i>et al.</i>	(NEAS)
WEHMANN	69	PR 178 2095	A.A. Wehmann <i>et al.</i>	(HARV, CASE, SLAC+)
ARMENISE	68	NC 54A 999	N. Armenise <i>et al.</i>	(BARI, BGNA, FIRZ+)
BATON	68	PR 176 1574	J.P. Baton, G. Laurens	(SACL)
CHUNG	68	PR 165 1491	S.U. Chung <i>et al.</i>	(LRL)
FOSTER	68	NP B6 107	M. Foster <i>et al.</i>	(CERN, CDEF)
GOUNARIS	68	PRL 21 244	G.J. Gounaris, J.J. Sakurai	
HUSON	68	PL 28B 208	R. Huson <i>et al.</i>	(ORSAY, MILA, UCLA)
HYAMS	68	NP B7 1	B.D. Hyams <i>et al.</i>	(CERN, MPIM)
LANZEROTTI	68	PR 166 1365	L.J. Lanzerotti <i>et al.</i>	(HARV)
PISUT	68	NP B6 325	J. Pisut, M. Roos	(CERN)
ASBURY	67B	PRL 19 865	J.G. Asbury <i>et al.</i>	(DESY, COLU)
BACON	67	PR 157 1263	T.C. Bacon <i>et al.</i>	(BNL)
EISNER	67	PR 164 1699	R.L. Eisner <i>et al.</i>	(PURD)
HUWE	67	PL 24B 252	D.O. Huwe <i>et al.</i>	(COLU)
HYAMS	67	PL 24B 634	B.D. Hyams <i>et al.</i>	(CERN, MPIM)
MILLER	67B	PR 153 1423	D.H. Miller <i>et al.</i>	(PURD)
ALFF-...	66	PR 145 1072	C. Alff-Steinberger <i>et al.</i>	(COLU, RUTG)



FERBEL	66	PL 21 111	T. Ferbel	(ROCH)
HAGOPIAN	66	PR 145 1128	V. Hagopian <i>et al.</i>	(PENN, SACL)
HAGOPIAN	66B	PR 152 1183	V. Hagopian, Y.L. Pan	(PENN, LRL)
JACOBS	66B	UCRL 16877	L.D. Jacobs	(LRL)
JAMES	66	PR 142 896	F.E. James, H.L. Kraybill	(YALE, BNL)
ROSS	66	PR 149 1172	M. Ross, L. Stodolsky	
SOEDING	66	PL B19 702	P. Soeding	
WEST	66	PR 149 1089	E. West <i>et al.</i>	(WISC)
BLIEDEN	65	PL 19 444	H.R. Blieden <i>et al.</i>	(CERN MMS Collab.)
CARMONY	64	PRL 12 254	D.D. Carmony <i>et al.</i>	(UCB)
GOLDHABER	64	PRL 12 336	G. Goldhaber <i>et al.</i>	(LRL, UCB)
ABOLINS	63	PRL 11 381	M.A. Abolins <i>et al.</i>	(UCSD)

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